specific atomic volume or volume per atom, $v_{0}$, (in cubic $\AA$ Angstroms, $\AA^{3}$ ) and distance between nearest neighbours $d_{0}$ (in $\AA$ ).

For a face centred cubic lattice, with the length of a unit cell $=a_{0}$ and 4 atoms/cell, the relationship between $a_{0}, d_{0}$ and $v_{0}$ is as follows:

$$
v_{0}=\frac{1}{d} \cdot a_{0}{ }^{3} \text { and } d_{0}{ }^{3}=(\sqrt{2}) \cdot v_{0}
$$

and

$$
a_{0}=1.4142 \cdot d_{0} \text { or } d_{0}=0.70711 \cdot a_{0}=1.1224 \cdot v_{0}{ }^{\frac{7}{2}} .
$$

For a hexagonal close packed lattice $d_{0}$ also equals $1 \cdot 1224 \cdot v_{0}{ }^{\frac{}{}{ }^{\text {. }} \text {. Furthermore, if }}$ $N_{\text {Avog. }}$ is Avogadro's number, $=60.225 \times 10^{22}$ atoms per g. atom,

$$
V_{\mathrm{at} .}=N_{\mathrm{Avog} .} \cdot v_{0}=\frac{\mathrm{at} . \mathrm{wt} .}{D_{0}}
$$

Throughout this paper we understand, by the compressibility $\kappa$, the isothermal compressibility proper, defined as ${ }^{(16)}$

$$
-\frac{1}{v_{0}}\left(\frac{d v}{d p}\right)_{T}
$$

and not the instantaneous compressibility

$$
-\frac{1}{v}\left(\frac{d v}{d p}\right)_{T}
$$

which has a greater tendency to increase at high pressure, because the factor $1 / v$ increases with pressure, while $1 / v_{0}$ is, of course, constant. We express $\kappa$ in c.g.s. units, i.e., in megabars ${ }^{-1}$, and not in $\left(\mathrm{kg} / \mathrm{cm}^{2}\right)^{-1}$, as has been Bridgman's preference, nor in atmospheres. The procedure adopted has the advantage that the data can be easily converted to other c.g.s.-units. The conversion factors used were:

$$
\begin{aligned}
1 \mathrm{~atm} & =1.01325 \times 10^{6} \text { dynes } / \mathrm{cm}^{2}\left(\text { or megabar }{ }^{-1}\right) \\
1 \mathrm{~kg} / \mathrm{cm}^{2} & =0.980665 \times 10^{6} \text { dynes } / \mathrm{cm}^{2}\left(\text { or megabar }{ }^{-1}\right)
\end{aligned}
$$

The data of Table 1 for Ne and $A r$ are based on Henshaw's ${ }^{(13)} a_{0}$ values obtained by neutron diffraction at $4 \cdot 2^{\circ} \mathrm{K}$. A slight correction was applied (same as in the case of Kr and Xe ) to extrapolate his values to $0^{\circ} \mathrm{K}$. The data for Kr and Xe are based on Smith's ${ }^{(14,15)}$ values of the density in the range of $20^{\circ}$ to $120^{\circ} \mathrm{K}$, obtained by X-ray diffraction. From the known coefficients of expansion, $\alpha$, in that range, and the fact that $\alpha=0$ at $0^{\circ} \mathrm{K}$, extrapolations were made to $0^{\circ} \mathrm{K}$. The data for emanation were estimated by us on the basis, originally developed by F. Paneth, that the differences in atomic volume of elements from period to period in the last periods of the Periodic System are practically constant.

Our values in Table 1 are close to the earlier ones, given in Table 2 and used by Bernardes ${ }^{(11)}$ in developing the theory of solid noble gases at $0^{\circ} \mathrm{K}$.
 pressure check StEWART's ${ }^{(7)}$ experimental data on Ne at $4 \cdot 2^{\circ} \mathrm{K}$ and his extrapolated data on Ar at $65^{\circ}$ and $77^{\circ} \mathrm{K}$ and on krypton at $77^{\circ} \mathrm{K}$.

[^0] (1949).


[^0]:    ${ }^{(16)}$ P. W. Bridgman, The Physics of High Pressure (1st. Ed) p. 169, Bell, London (1931); reprinted

